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Discrete random walk models for space-time fractional diffusion

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Abstract

A physical-mathematical approach to anomalous diffusion may be based on generalized diffusion equations (containing derivatives of fractional order in space or/and time) and related random walk models. By space-time fractional diffusion equation we mean an evolution equation obtained from the standard linear diffusion equation by replacing the second-order space derivative with a Riesz-Feller derivative of order $\alpha \in (0, 2]$ and skewness $\theta (|\theta| \leq \min\{\alpha, 2 - \alpha\})$, and the first-order time derivative with a Caputo derivative of order $\beta \in (0, 1]$. Such evolution equation implies for the flux a fractional Fick's law which accounts for spatial and temporal non-locality. The fundamental solution (for the Cauchy problem) of the fractional diffusion equation can be interpreted as a probability density evolving in time of a peculiar self-similar stochastic process that we view as a generalized diffusion process. By adopting appropriate finite-difference schemes of solution, we generate models of random walk discrete in space and time suitable for simulating random variables whose spatial probability density evolves in time according to this fractional diffusion equation. © 2002 Elsevier Science B.V. All rights reserved.

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1. Introduction

It is well known that the fundamental solution (or *Green function*) for the Cauchy problem of the linear diffusion equation can be interpreted as a Gaussian (normal) probability density function (pdf) in space, evolving in time. All the moments of this pdf are finite; in particular, its variance is proportional to the first power of time, a note-worthy property of the *standard diffusion* that can be understood by means of an unbiased random walk model for the *Brownian motion*.

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In recent years a number of master equations have been proposed for random walk models that turn out to be beyond the classical Brownian motion, see e.g. [32]. In particular, evolution equations containing fractional derivatives have gained revived interest in that they are expected to provide suitable mathematical models for describing phenomena of anomalous diffusion and transport dynamics in complex systems, see e.g. [1,10,15,28–30,34–36,40,48,49,52,57,61–64]. For a recent review, we refer to Metzler and Klafter [41] where other references are found.

Here, we intend to present our original approach to the topic that, being not considered in [41], could offer some novel and inspiring inspections to the phenomenon of anomalous diffusion which is of great interest in chemical physics. In this paper, we complement and revisit some of our previous results found e.g. in [19,23,26]. We first show that our proposed fractional diffusion equations can be derived from generalized Fick's laws which account for spatial and/or temporal non-locality. Then we pay attention to the fact that the fundamental solutions (or Green functions) of our diffusion equations provide spatial probability densities evolving in time, related to *self-similar* stochastic processes, that we view as generalized (or *fractional*) diffusion processes to be properly understood through suitable random walk models. More precisely, we replace the second-order space derivative or/and the first-order time derivative by a suitable integrodifferential operator, which can be interpreted as a space or time derivative of fractional order $\alpha \in$ (0,2] or $\beta \in (0,1]$, respectively.¹ The space fractional derivative is required to depend also on a real parameter θ (the *skewness*) subjected to the restriction $|\theta| \leq \min\{\alpha, 2 - \alpha\}$. Correspondingly, the generalized equation will be referred to as the strictly space fractional diffusion equation of order α and skewness θ if $\alpha \in (0, 2)$ and $\beta = 1$, or the *strictly time fractional diffusion* equation of order β if $\alpha = 2$ and $\beta \in (0, 1)$. In general, allowing a $\alpha \in (0, 2)$ and

 $\beta \in (0, 1)$, we have the strictly space–time fractional diffusion equation of order α , β and skewness θ . Of course, in the case { $\alpha = 2, \beta = 1$ } we recover the *standard diffusion* which leads to the Gaussian probability density and to the classical *Brownian motion*.

For the strictly space fractional diffusion of order α ({ $0 < \alpha < 2, \beta = 1$ }) we generate the class of (non-Gaussian) Lévy stable densities of index α and skewness θ ($|\theta| \leq \min\{\alpha, 2 - \alpha\}$), according to the Feller parameterization. As known, these densities exhibit fat tails with an algebraic decay $\propto |x|^{-(\alpha+1)}$. We thus obtain a special class of *Markovian* stochastic processes, called *stable Lévy motions*, which exhibit infinite variance associated to the possibility of arbitrarily large jumps (*Lévy flights*).

For the strictly time fractional diffusion of order β ({ $\alpha = 2, 0 < \beta < 1$ }) we generate a class of symmetric densities whose moments of order 2*n* are proportional to the *n* β power of time. We thus obtain a class of *non-Markovian* stochastic processes (they possesses a memory!) which exhibit a variance consistent with *slow anomalous diffusion*.

For the *strictly space-time fractional diffusion* of (composite) order $\alpha \in (0, 2), \beta \in (0, 1)$ we generate a class of densities (symmetric or not symmetric according to $\theta = 0$ or $\theta \neq 0$) which exhibit fat tails with an algebraic decay $\propto |x|^{-(\alpha+1)}$. Thus they belong to the domain of attraction of the Lévy stable densities of index α and can be called *fractional* stable densities. The related stochastic processes, by possessing the characteristics of the previous two classes, are non-Markovian and exhibit infinite variance; however, the possibility of arbitrary large jumps is contrasted by memory effects. Furthermore, we mention the cases $\alpha = \beta$ for which it is possible to derive the Green function in closed analytical form: we refer to these cases as to neutral diffusion.

We shall prove that in any case the corresponding Green function can be interpreted as a spatial probability density evolving in time with a self-similarity property having scaling exponent $v = \beta/\alpha$. This allows us to limit ourselves to consider the expression of the Green function at a fixed time, say t = 1, namely to the so-called *reduced Green function*. To approximate the time

¹ We remind that the term "fractional" is a misnomer since the order can be a real number and thus is not restricted to be rational. The term is kept only for historical reasons, see e.g. [22]. Our fractional derivatives are required to coincide with the standard derivatives of integer order as soon as $\alpha = 2$ (not as $\alpha = 1!$) and $\beta = 1$.

evolution of all the above densities we propose finite difference schemes, discrete in space and time, for the fractional derivatives. By taking care in constructing these schemes, namely by requiring them to be conservative and non-negativity preserving, they can be interpreted as discrete random walk models for simulating particle paths by the Monte Carlo technique. By properly scaled transition to vanishing space and time steps, these models can be shown to converge to the corresponding continuous processes.²

The paper is divided as follows. In Section 2, we first present our space and time fractional diffusion equations providing the definitions of the space and time fractional derivatives based on their Fourier and Laplace representations, respectively. Then we show how to derive them from generalized Fick's laws. Section 3 is devoted to the Green functions, pointing out their similarity properties. We provide the representations of the corresponding reduced Green functions in terms of Mellin-Barnes integrals which allow us to obtain their computational expressions. In Section 4, we first discuss the discrete random walk approach to the Brownian motion, which is based on the wellknown discretization of the second-order space derivative and the first-order time derivative entering the standard diffusion equation. Then, by properly discretizing the space-fractional derivative we generalize the above approach to the more general Markovian case of strictly space fractional *diffusion*. Section 5 is devoted to the extension of the above approach to the non-Markovian cases of strictly time fractional diffusion and strictly spacetime fractional diffusion equations. Section 6 is devoted to the numerical results of our random walks produced in some case-studies and to the concluding discussions. For possible convenience of the reader we have reserved Appendices A and B for treating with some detail the Riesz-Feller and Caputo fractional derivatives, respectively.

2. The space-time fractional diffusion equation

By replacing in the standard diffusion equation

$$\frac{\partial}{\partial t}u(x,t) = \frac{\partial^2}{\partial x^2}u(x,t), -\infty < x < +\infty, \ t \ge 0,$$
(2.1)

where u = u(x,t) is the (real) field variable, the second-order space derivative and the first-order time derivative by suitable *integro-differential* operators, which can be interpreted as a space and time derivative of fractional order we obtain a generalized diffusion equation which may be referred to as the *space-time fractional* diffusion equation. We write this equation as

$${}_{t}D_{*}^{\beta}u(x,t) = {}_{x}D_{\theta}^{\alpha}u(x,t),$$

- $\infty < x < +\infty, t \ge 0,$ (2.2)

where α, θ, β are real parameters restricted to

$$0 < \alpha \leq 2, \quad |\theta| \leq \min\{\alpha, 2 - \alpha\}, \quad 0 < \beta \leq 1.$$

In (2.2) ${}_{x}D^{\alpha}_{\theta}$ is the *Riesz–Feller fractional derivative* (in space) of order α and skewness θ , and ${}_{t}D^{\beta}_{*}$ is the *Caputo fractional derivative* (in time) of order β .

The definitions of these fractional derivatives are more easily understood if given in terms of Fourier and Laplace transforms, respectively.

In terms of the Fourier transform the *Riesz– Feller fractional derivative in* space is defined as

$$\mathscr{F}_{\{x} D^{\alpha}_{\theta} f(x); \kappa\} = -\psi^{\alpha}_{\theta}(\kappa) \hat{f}(\kappa), \psi^{\alpha}_{\theta}(\kappa) = |\kappa|^{\alpha} \mathrm{e}^{\mathrm{i}(\mathrm{sign}\,\kappa)\theta\pi/2},$$
(2.4)

where $\hat{f}(\kappa) = \mathscr{F}\{f(x);\kappa\} = \int_{-\infty}^{+\infty} e^{i\kappa x} f(x) dx$. In other words $_{x}D_{\theta}^{z}$ is a pseudo-differential operator ³ with $_{x}\widehat{D_{\theta}^{z}}(\kappa) = -\psi_{\theta}^{z}(\kappa)$, which is the logarithm of the characteristic function of the generic *strictly stable*

 $^{^{2}}$ This was shown by Gorenflo and Mainardi for the space fractional diffusion in [23–25]. For the general case it will be shown in a next paper.

³ Let us recall that a generic pseudo-differential operator A, acting with respect to the variable $x \in \mathbf{R}$, is defined through its Fourier representation, namely $\widehat{Af}(\kappa) = \widehat{A}(\kappa)\widehat{f}(\kappa)$, where $\widehat{A}(\kappa)$ is referred to as the symbol of A. The *n*th derivative operator ${}_{x}D^{n} = d^{n}/dx^{n}$ is a special case with symbol $\widehat{xD^{n}}(\kappa) = (-i\kappa)^{n}$. Generally speaking, a pseudo-differential operator A turns out to be defined through a kernel of a space convolution integral; this kernel is thus a sufficiently well-behaved function (absolutely) integrable in \mathbf{R} which degenerates to a delta-type distribution when $A = {}_{x}D^{n}$. Furthermore, as a matter of fact, the symbol is given by the rule $\widehat{A}(\kappa) = (Ae^{-i\kappa x})e^{+i\kappa x}$.



Fig. 1. The Feller–Takayasu diamond of the generic Riesz– Feller derivative.

(in the Lévy sense) probability density, according to the Feller parameterization [12,13].

We note that the allowed region for the parameters α and θ turns out to be a diamond in the plane $\{\alpha, \theta\}$ with vertices in the points (0,0), (1,1), (2,0), (1,-1), that we call the *Feller–Takayasu diamond*⁴, see Fig. 1.

For $\alpha = 2$ (henceforth $\theta = 0$) we ${}_{x}D_{0}^{2}(\kappa) = -\kappa^{2} = (-i\kappa)^{2}$, so we recover the standard second derivative. More generally for $\theta = 0$ we have ${}_{x}\overline{D}_{0}^{\alpha}(\kappa) = -|\kappa|^{\alpha}) = -(\kappa^{2})^{\alpha/2}$ so

$$_{x}D_{0}^{\alpha} = -\left(-\frac{\mathrm{d}^{2}}{\mathrm{d}x^{2}}\right)^{\alpha/2} = _{x}D^{2}.$$
 (2.5)

In this case we refer to the LHS of (2.5) as simply to the *Riesz fractional derivative* of order α . We refer to Appendix A for the explicit expression of the generic Riesz–Feller derivative.

Let us now consider the *Caputo fractional derivative* in time. Following the original idea by Caputo [3,4], see also [7,8,22,50], a proper time fractional derivative of order $\beta \in (0, 1)$, useful for physical applications, may be defined in terms of the following rule for the Laplace transform:

$$\mathcal{L}\{{}_{t}D^{\beta}_{*}f(t);s\} = s^{\beta}\hat{f}(s) - s^{\beta-1}f(0^{+}), 0 < \beta < 1,$$
(2.6)

where $\tilde{f}(s) = \mathscr{L}{f(t); s} = \int_0^\infty e^{-st} f(t) dt$. Then the *Caputo fractional derivative* of f(t) turns out to be defined as

$${}_{t}D_{*}^{\beta}f(t) := \begin{cases} \frac{1}{\Gamma(1-\beta)} \int_{0}^{t} \frac{f^{(1)}(\tau)}{(t-\tau)^{\beta}} \, \mathrm{d}\tau, & 0 < \beta < 1, \\ \frac{\mathrm{d}f(t)}{\mathrm{d}t}, & \beta = 1. \end{cases}$$
(2.7)

In other words the ${}_{t}D_{*}^{\beta}$ is required to generalize the well-known rule for the Laplace transform of the first derivative of a given (causal) function keeping the standard initial value of the function itself. We refer to Appendix B for the relation of this derivative to the more common Riemann Liouville fractional integral and derivative. Here, we report the most relevant formula (B.6) which provides alternative expressions for the Caputo fractional derivative for $0 < \beta < 1$:

$${}_{t}D_{*}^{\beta}f(t) = \begin{cases} \frac{1}{\Gamma(1-\beta)} \frac{d}{dt} \int_{0}^{t} \frac{f(\tau) - f(0^{+})}{(t-\tau)^{\beta}} d\tau, \\ \frac{1}{\Gamma(1-\beta)} \frac{d}{dt} \int_{0}^{t} \frac{f(\tau)}{(t-\tau)^{\beta}} d\tau - f(0^{+}) \frac{t^{-\beta}}{\Gamma(1-\beta)}. \end{cases}$$

$$(2.8)$$

It is worth to note that the time fractional derivative in the LHS of Eq. (2.2) can be removed by a suitable fractional integration, see (B.4), leading to the alternative form

$$u(x,t) = u(x,0^{+}) + \frac{1}{\Gamma(\beta)} \int_{0}^{t} {}_{x} D_{\theta}^{x} u(x,\tau) \frac{d\tau}{(t-\tau)^{1-\beta}}.$$
(2.9)

Differentiating with respect to time we have another equivalent form

$$\frac{\partial}{\partial t}u(x,t) = \frac{1}{\Gamma(\beta)} \frac{\partial}{\partial t} \left\{ \int_0^t {}_x D^x_{\theta} u(x,\tau) \frac{\mathrm{d}\tau}{\left(t-\tau\right)^{1-\beta}} \right\}.$$
(2.10)

It is well known that in the case of standard diffusion Eq. (2.1) can be derived from the *continuity equation*:

$$\frac{\partial}{\partial t}u(x,t) + \frac{\partial}{\partial x}F[u(x,t)] = 0, \qquad (2.11)$$

where F is the *flux* given by

⁴ Our notation for the stable distributions has been adapted from the original one by Feller. From 1998, see [23], we have found it as the most convenient among the others available in the literature, see e.g. [31,33,43,44,54,55,62,67]. Furthermore, this notation has the advantage that all the class of the *strictly stable* densities are represented. As far as we know, the diamond representation in the plane $\{\alpha, \theta\}$ was formerly given by Takayasu in his 1990 book on *Fractals* [59].

$$F[u(x,t)] = -\frac{\partial}{\partial x}u(x,t).$$
(2.12)

Whereas Eq. (2.11) is related to a conservation law of a physical quantity (therefore it has a universal character), Eq. (2.12) is a phenomenological law which states that the flux is proportional to the gradient of the field variable (the transported quantity) with opposite sign. It is met in several physical contexts with different names: e.g. when u is the temperature it is known as Fourier's law, when u is the pore pressure as Darcy's law, when u is a concentration of particles, as Fick's law. Here, we use the last terminology in view of the possible applications in chemical physics. We recall that Fick's law is essentially an empirical law, which represents the simplest (local in space and time) relationship between the flux F and the gradient of the concentration *u* observable in several physical phenomena. As a matter of fact, for some experimental evidences in complex transport phenomena, this law can be replaced by a more suitable phenomenological relationship which may account for possible non-local and memory effects, without violating the conservation law expressed by the continuity equation (2.11). Now it is not difficult to derive our fractional diffusion equation in the form (2.10) from a generalized Fick's law in which a suitable space-time operator depending on α, θ and β is acting on the gradient. After simple manipulations based on recent results by Paradisi et al. [48,49] and Gorenflo et al. [26], we can write

$$F(x,t) = {}_{t}D^{1-\beta}{}_{x}P^{\alpha}_{\theta} \left[-\frac{\partial}{\partial x}u(x,t) \right], \qquad (2.13)$$

where ${}_{t}D^{1-\beta}$ denotes the Riemann–Liouville fractional derivative of order $1 - \beta$ (in time), see (B.1), and ${}_{x}P^{\alpha}_{\theta}$ is the pseudo-differential operator with symbol

$$\widehat{{}_{x}P_{\theta}^{\alpha}}(\kappa) := \frac{\widehat{{}_{x}D_{\theta}^{\alpha}(\kappa)}}{\widehat{{}_{x}D^{2}}(\kappa)} = |\kappa|^{\alpha-2} \mathrm{e}^{\mathrm{i}(\mathrm{sign}\,\kappa)\theta\pi/2}.$$
(2.14)

For $\alpha = 2$ and $\beta = 1$ we recover the standard Fick's law since in this case ${}_{t}D^{0}{}_{x}P_{0}^{2} = I$ (Identity). For $1 < \alpha < 2$ Eq. (2.13) results to be a non-local connection between the flux and the concentration gradient both from temporal and spatial view points. For the nature of the operators involved, Eq. (2.13) can be referred to as the *fractional Fick's law:* it turns out to be consistent with the spacetime fractional diffusion equation (2.2). We note that for $0 < \alpha \le 1$ Eq. (2.13) is meaningless since the symbol of the pseudo-differential P_{θ}^{α} in (2.14) exhibits at $\kappa = 0$ a singularity not Fourier integrable, which means that the kernel is not integrable in **R**.⁵ Generalized Fick's laws with fractional derivatives have also been considered by other authors, including e.g. [5,6,65].

3. The Green function for the space-time fractional diffusion

The fundamental solution (or the *Green function*) for the space–time fractional diffusion is intended to be the solution of the governing equation (2.2), or (2.9) or (2.10), corresponding to the initial condition $u(x, 0^+) = \delta(x)$, It will be denoted by $G^{\theta}_{\alpha,\beta}(x,t)$. In the case of standard diffusion, see Eq. (2.1), the Green function is nothing but the Gaussian probability density function with variance $\sigma^2 = 2t$, namely

$$G_{2,1}^{0}(x,t) = \frac{1}{2\sqrt{\pi}} t^{-1/2} e^{-x^{2}/(4t)}.$$
(3.1)

In the general case, following the arguments by Mainardi et al. [37] we can prove that $G^{\theta}_{\alpha,\beta}(x,t)$ is still a probability density evolving in time with the noteworthy scaling property

$$G^{\theta}_{\alpha,\beta}(x,t) = t^{-\beta/\alpha} K^{\theta}_{\alpha,\beta}(x/t^{\beta/\alpha}).$$
(3.2)

$$F(x,t) = {}_t D^{1-\beta} {}_x Q^{\alpha}_{\theta} \ [u(x,t)],$$

where ${}_{x}Q^{\alpha}_{\theta}$ is the pseudo-differential operator with symbol

$$\widehat{{}_{x}\mathcal{Q}_{\theta}^{\alpha}}(\kappa) := {}_{x}P_{\theta}^{\alpha}(\kappa) \widehat{{}_{x}D^{1}}(\kappa) = |\kappa|^{\alpha-1} \mathrm{e}^{\mathrm{i}(\mathrm{sign}\,\kappa)(\theta+1)\pi/2}.$$

Physical reasons, however, lead us to avoid the range $0 < \alpha \le 1$. In fact, for this range, $\widehat{\Omega}_{\delta}^{\alpha}(\kappa)$ would be a decreasing (or constant) function of $|\kappa|$, which means that the contribution to the flux of the larger scales would be greater than (or equal to) that of the smaller ones, which is meaningless.

⁵ From a purely mathematical view point one could overcome the above trouble for $0 < \alpha \le 1$ by stating the relationship between the flux *F* and the concentration *u* as

Here, $x/t^{\beta/\alpha}$ acts as the similarity variable and $K^{\theta}_{\alpha,\beta}(\cdot)$ as the *reduced Green function*. For the analytical and computational determination of the reduced Green function, from now on we restrict our attention to x > 0 because of the *symmetry relation* $K^{\theta}_{\alpha,\beta}(-x) = K^{-\theta}_{\alpha,\beta}(x)$. Mainardi et al. [37] have provided (for x > 0) the Mellin–Barnes integral representation⁶

$$K^{\theta}_{\alpha,\beta}(x) = \frac{1}{\alpha x} \frac{1}{2\pi i} \int_{\gamma-i\infty}^{\gamma+i\infty} \frac{\Gamma(\frac{s}{\alpha})\Gamma(1-\frac{s}{\alpha})\Gamma(1-s)}{\Gamma(1-\frac{\beta}{\alpha}s)\Gamma(\rho s)\Gamma(1-\rho s)} x^{s} ds,$$

$$\rho = \frac{\alpha-\theta}{2\alpha}, \qquad (3.3)$$

where $0 < \gamma < \min\{\alpha, 1\}$.

We recognize from the footnote 6 that Eq. (3.3) by changing s into -s can be interpreted as a Mellin transform pair that allows us to write the Mellin transform of $xK^{\theta}_{\alpha,\beta}(x)$ as

$$\int_{0}^{+\infty} K^{\theta}_{\alpha,\beta}(x) x^{s} \, \mathrm{d}x = \frac{1}{\alpha} \frac{\Gamma(-s/\alpha)\Gamma(1+s/\alpha)\Gamma(1+s)}{\Gamma(1+\beta s/\alpha)\Gamma(-\rho s)\Gamma(1+\rho s)}, \\ -\min\{\alpha,1\} < \Re(s) < 0.$$
(3.4)

In order to include s = 0 in the convergence strip (so, in particular, the integral of $K^{\theta}_{\alpha,\beta}(x)$ in \mathbf{R}^+_0 can be evaluated) we properly use in (3.4) the functional equation $\Gamma(1+z) = z\Gamma(z)$ to obtain

$$\begin{split} &\int_{0}^{+\infty} K^{\theta}_{\alpha,\beta}(x) x^{s} \, \mathrm{d}x = \rho \, \frac{\Gamma(1-s/\alpha) \Gamma(1+s/\alpha) \Gamma(1+s)}{\Gamma(1-\rho s) \Gamma(1+\rho s) \Gamma(1+\beta s/\alpha)}, \\ &- \min\{\alpha,1\} < \Re(s) < \alpha. \end{split}$$

⁶ The names refer to the two authors, who in the first 1910s developed the theory of these integrals using them for a complete integration of the hypergeometric differential equation. However, as pointed out in [11] (Vol. 1, Chapter 1, Section 1.19, p. 49), these integrals were first used by Pincherle in 1888. For a revisited analysis of the pioneering work of Pincherle (Professor of Mathematics at the University of Bologna from 1880 to 1928) we refer to the recent paper by Mainardi and Pagnini [38]. As a matter of fact this type of integral turns out to be useful in inverting the Mellin transform, as shown hereafter. If

$$\mathscr{M}{f(r);s} = f^*(s) = \int_0^{+\infty} f(r)r^{s-1} dr, \quad \gamma_1 < \Re(s) < \gamma_2$$

denotes the Mellin transform of a sufficiently well-behaved function f(r), the inversion is provided by

$$\mathcal{M}^{-1}\{f^*(s);r\} = f(r) = \frac{1}{2\pi i} \int_{\gamma-i\infty}^{\gamma+i\infty} f^*(s)r^{-s} ds,$$

where $r > 0, \ \gamma = \Re(s), \ \gamma_1 < \gamma < \gamma_2.$

In particular, we find $\int_0^{+\infty} K_{\alpha,\beta}^{\theta}(x) \, dx = \rho$ (with $\rho = 1/2$ if $\theta = 0$).

We note that Eq. (3.5) is strictly valid as soon as cancellations in the "gamma fraction" at the RHS are not possible. Then this equation allows us to evaluate (in \mathbf{R}_{0}^{+}) the (absolute) moments of order δ for the Green function such that $-\min\{\alpha,1\} < \delta < \alpha$. In other words, it states that $K_{\alpha,\beta}^{\theta}(x) = \mathcal{O}(x^{-(\alpha+1)})$ as $x \to +\infty$. When cancellations occur in the "gamma fraction" the range of δ may change. An interesting case is $\{\alpha = 2, \theta = 0, 0 < \beta \leq 1\}$ (*time-fractional diffusion* including *standard diffusion*), where Eq. (3.5) reduces to

$$\int_{0}^{+\infty} K_{2,\beta}^{0}(x) x^{s} dx = \frac{1}{2} \frac{\Gamma(1+s)}{\Gamma(1+\beta s/2)},$$

$$\Re(s) > -1.$$
(3.6)

This result proves the existence of all moments of order $\delta > -1$ for the corresponding Green function. In virtue of (3.2), (3.6) we have

$$\int_{-\infty}^{+\infty} |x|^{\delta} G_{2,\beta}^{0}(x,t) \, \mathrm{d}x = \frac{\Gamma(\delta+1)}{\Gamma(\beta\delta/2+1)} t^{\beta\delta/2},$$

$$\delta > -1, \quad 0 < \beta \le 1, \tag{3.7}$$

and, for $\delta = 2$, the following formula for the variance:

$$\int_{-\infty}^{+\infty} x^2 G_{2,\beta}^0(x,t) \, \mathrm{d}x = \frac{2}{\Gamma(\beta+1)} t^{\beta},$$
$$0 < \beta \leqslant 1. \tag{3.8}$$

We note that the Mellin–Barnes integral representation allows us to construct computationally the fundamental solutions of Eq. (2.1) for any triplet { α , β , θ } by matching their convergent and asymptotic expansions, see [37,47]. Readers acquainted with Fox *H* functions can recognize in (3.3) the representation of a certain function of this class, see e.g. [29,39,40,53,56–58,62]. Unfortunately, as far as we know, computing routines for this general class of special functions are not yet available.

Let us now point out the main characteristics of the peculiar cases of *strictly space fractional diffusion, strictly time fractional diffusion*, and *neutral fractional diffusion* based on the results stated in [37]. For $\beta = 1$ and $0 < \alpha < 2$ (strictly space fractional diffusion) we recover the class of the strictly stable (non-Gaussian) densities exhibiting heavy tails (with the algebraic decay $\propto |x|^{-(\alpha+1)}$) and infinite variance. For $\alpha = 2$ and $0 < \beta < 1$ (strictly time fractional diffusion) we recover the class of the Wright-type densities exhibiting stretched exponential tails and finite variance proportional to t^{β} . Mathematical details on these two classes of probability densities can be found in [37]; for further reading we refer to Schneider [56] for stable densities, and to Gorenflo et al. [20,21] for the Wright-type densities.

As for the stochastic processes governed by these distributions we can expect the following.

For the case of non-Gaussian stable densities we expect a special class of Markovian processes, called stable Lévy motions, which exhibit infinite variance associated to the possibility of arbitrarily large jumps (*Lévy flights*), whereas for the case of Wright-type densities we expect a class of stochastic non-Markovian processes, which exhibit a (finite) variance consistent with slow anomalous diffusion.

For the special case $\alpha = \beta \leq 1$ (*neutral diffusion*) we obtain from (3.3) an elementary (non-negative) expression

$$\begin{split} K^{\theta}_{\alpha,\alpha}(x) &= \frac{1}{\alpha x} \frac{1}{2\pi i} \int_{\gamma - i\infty}^{\gamma + i\infty} \frac{\Gamma\left(\frac{s}{\alpha}\right) \Gamma(1 - \frac{s}{\alpha})}{\Gamma(\rho s) \Gamma(1 - \rho s)} x^{s} \, \mathrm{d}s \\ &= \frac{1}{\alpha x} \frac{1}{2\pi i} \int_{\gamma - i\infty}^{\gamma + i\infty} \frac{\sin(\pi \rho s)}{\sin(\pi s/\alpha)} x^{s} \, \mathrm{d}s \\ &= \frac{1}{\pi} \frac{x^{\alpha - 1} \sin[\frac{\pi}{2}(\alpha - \theta)]}{1 + 2x^{\alpha} \cos[\frac{\pi}{2}(\alpha - \theta)] + x^{2\alpha}}, \\ x &> 0, \end{split}$$
(3.9)

where $0 < \gamma < \alpha$.

For the generic case of *strictly space-time diffusion* $(0 < \alpha < 2, 0 < \beta < 1)$, including neutral diffusion for $\alpha = \beta < 1$, Mainardi et al. [37] have proven the non-negativity of the corresponding reduced Green function and consequently its interpretation as probability density. In this case we obtain a class of probability densities (symmetric or non-symmetric according to $\theta = 0$ or $\theta \neq 0$), which exhibit heavy tails with an algebraic decay $\propto |x|^{-(\alpha+1)}$. Thus they belong to the domain of at-

traction of the Lévy stable densities of index α and can be referred to as *fractional stable densities*, according to a terminology proposed by Uchaikin [60]. The related stochastic processes are expected to possess the characteristics of the previous two classes. Indeed, they are non-Markovian being $\beta < 1$ and exhibit infinite variance associated to the possibility of arbitrarily large jumps (being $\alpha < 2$).

4. The discrete random walk models for the Markovian fractional diffusion

It is known that a numerical approach to the standard diffusion equation (2.1) based on a proper finite difference scheme provides a discrete Markovian random walk model for the classical Brownian motion, see e.g. [66].

In this section we intend to generalize this approach (that will be hereafter recalled) in order to provide a discrete Markovian random walk model for the Lévy stable motion of any order $\alpha \in (0, 2)$ and skewness θ restricted as in (2.3). For this purpose we present a notable finite-difference approach to the strictly space fractional diffusion equation subjected to relevant restrictions, as we shall show in the following.

The common starting point of our analysis is obviously the discretization of the space-time domain by grid points and time instants as follows:

$$\begin{cases} x_j = jh, \quad h > 0, \quad j = 0, \pm 1, \pm 2, \dots, \\ t_n = n\tau, \quad \tau > 0, \quad n = 0, 1, 2, \dots, \end{cases}$$
(4.1)

where the steps h and τ are assumed to be small enough. The dependent variable u is then discretized by introducing $y_i(t_n)$ as

$$y_j(t_n) \approx \int_{x_j-h/2}^{x_j+h/2} u(x,t_n) \,\mathrm{d}x \approx h \,\, u(x_j,t_n).$$
 (4.2)

4.1. The discrete random walk model for the standard diffusion

Let us now consider the standard diffusion equation (2.1). With the quantities $y_j(t_n)$ so intended, we replace Eq. (2.1), after multiplication by the spatial mesh-width *h*, by the finite-difference equation

$$\frac{y_j(t_{n+1}) - y_j(t_n)}{\tau} = \frac{y_{j+1}(t_n) - 2y_j(t_n) + y_{j-1}(t_n)}{h^2} \quad (4.3)$$

accepting that for positive *n* in (4.3) we have approximate instead of exact equality. Since we are interested to approximate the fundamental solution (the Green function), we must equip (4.3) with the initial condition $y_j(0) = \delta_{j0}$, where the Kronecker symbol represents the discrete counterpart of the Dirac delta function.

This approach can be interpreted as a discrete (in space and time) *redistribution process* of some extensive quantity provided it is conservative. If the extensive quantity is non-negative, e.g. mass or a sojourn probability, we have to preserve its nonnegativity. In the first case, the $y_j(t_n)$ are imagined as clumps of mass, sitting at grid points $x = x_j$ in instants $t = t_n$, which collect approximatively the total mass in the interval $x_j - h/2 < x \le x_j + h/2$. In the second case, the $y_j(t_n)$ may be interpreted as the probability of sojourn in point x_j at time t_n for a particle making a *random walk* on the spatial grid in discrete instants. From now on, we agree to pursue the probabilistic point of view.

In order to have a *conservative* and *non-negativity preserving* redistribution process, the discrete variable y_i is subjected to the conditions

$$\sum_{j=-\infty}^{+\infty} y_j(t_n) = \sum_{j=-\infty}^{+\infty} y_j(0), \quad y_j(t_n) \ge 0$$

for all $j \in \mathbf{Z}, \ \mathbf{n} \in \mathbf{N}_0.$ (4.4)

We easily recognize that our discrete redistribution process is akin to a *Markov chain:* when time proceeds from $t = t_n$ to $t = t_{n+1}$, the sojourn-probabilities are redistributed according to the *transition law*

$$y_j(t_{n+1}) = \sum_{k=-\infty}^{\infty} p_k y_{j-k}(t_n), \quad j \in \mathbf{Z}, \ \mathbf{n} \in \mathbf{N_0}, \quad (4.5)$$

where the p_k denote suitable *transfer coefficients*, which represent the probability of transition from x_{j-k} to x_j (likewise from x_j to x_{j+k}). The transfer coefficients are to be found consistently with the finite-difference equation (4.3) equipped with the proper initial condition. The process turns out to be both spatially homogeneous (the probability p_k of jumping from a point x_j to a point x_{j+k} not depending on j) and time-stationary (the p_k not depending on n), as is advised when considering our Cauchy problem and the definition of the difference operators. Furthermore, the transfer coefficients must satisfy the conditions

$$\sum_{k=-\infty}^{\infty} p_k = 1, \quad p_k \ge 0, \quad k = 0, \pm 1, \pm 2, \dots$$
 (4.6)

The transfer coefficients in our special case are easily deduced from (4.3) and (4.5): they turn out to be

$$p_0 = 1 - 2\frac{\tau}{h^2}, \quad p_{\pm 1} = \frac{\tau}{h^2}, \quad p_{\pm k} = 0, \quad k = 2, 3, \dots$$

(4.7)

subject to the condition

$$0 < \mu := \frac{\tau}{h_2} \leqslant \mu_{\max} = \frac{1}{2}.$$
(4.8)

We refer to Eqs. (4.6)–(4.8) as the basic equations for the standard random walk model for the Gaussian process. The constant μ will be denoted as the *scaling parameter* of the standard diffusion equation. We recognize that only jumps of one step to the right or to the left or jumps of width zero occur. This corresponds to a well-known simple approximate realization of the *Brownian motion*.

The finite difference scheme (4.3) can be used for producing *sample paths* of individual particles performing the Brownian motion and for producing *histograms* of the approximate realization of the Gaussian density, by simulating many individual paths with the same number of time steps and making statistics of the final positions of the particles.

Our simulations, based on 10,000 realizations, have been limited to the interval $|x| \leq 5$, where we have considered the random walks to take place, have produced an histogram for the Gaussian density at t = 1, see Fig. 2 (left plate). Then, particles leaving this space have been ignored. In Fig. 2 (right plate) we have displayed a particular sample path or trajectory (up) and the corresponding increment series (below) obtained with N = 500 time steps, by taking an intermediate value for the scaling factor ($\mu = 0.4 < 1/2$) and reasonable values for the space and time steps (h = 0.0707, $\tau = 0.002$) in order to get $N\tau = t = 1$. Of course the histogram has been depicted by



Fig. 2. Histogram (left) and a sample path with increments (right) for standard diffusion { $\alpha = 2, \beta = 1, \theta = 0$ } (Brownian motion).

adopting a space step much larger than for the sample path, namely h = 0.25.

4.2. A discrete random walk model for the strictly space fractional diffusion

Discretizing all the variables as done for the standard diffusion equation, namely introducing a space–time mesh of widths h and τ and a discrete variable $y_j(t_n)$ interpreted as in (4.2), the *essential idea* is to replace the *strictly space fractional dif-fusion* equation by the finite-difference equation

$$\frac{y_j(t_{n+1} - y_j(t_n))}{\tau} = {}_h D^{\alpha}_{\theta} y_j(t_n),$$

$$0 < \alpha < 2, \quad |\theta| \le \min\{\alpha, 2 - \alpha\}$$
(4.9)

where the difference operator ${}_{h}D^{\alpha}_{\theta}$ is intended to converge to D^{α}_{θ} (discussed in Appendix A) as $h \to 0$. As usual, we have adopted a forward difference quotient in time at level $t = t_n$ for approximating the first-order time derivative. For ${}_{h}D^{\alpha}_{\theta}$ we require a scheme which must reduce as $\alpha = 2$ to a symmetric second-order difference quotient in space at level $t = t_n$, which has been adopted for approximating the second-order space derivative. Furthermore, the finite-difference equation (4.9) is required to be consistent with a *conservative*, *non-negativity preserving redistribution process*, subject to the initial condition $y_j(0) = \delta_{j0}$. Referring to Eqs. (A.4), (A.8) and (A.10), (A.11), we write

$${}_{h}D^{\alpha}_{\theta}y_{j}(t_{n}) = \begin{cases} -\left[c_{+}(\alpha,\theta)_{h}D^{\alpha}_{+} + c_{-}(\alpha,\theta)_{h}D^{\alpha}_{-}\right]y_{j}(t_{n}), \\ \alpha \neq 1, \\ \left[\cos(\theta\pi/2)_{h}D^{1}_{0} + \sin(\theta\pi/2)_{h}D_{\pm}\right]y_{j}(t_{n}), \\ \alpha = 1. \end{cases}$$
(4.10)

With the notation ${}_{h}D_{\pm}$ we intend to adopt ${}_{h}D_{+}$ when $0 < \theta \leq 1$ and ${}_{h}D_{-}$ when $-1 \leq \theta < 0$.

We thus must find suitable finite-difference schemes for the pseudo-differential operators entering the Feller derivative, namely ${}_{h}D^{\alpha}_{\pm}$ (Weyl fractional derivatives of order α) if $a \neq 1$ and ${}_{h}D^{1}_{0}$ (Riesz derivative of first-order) if $\alpha = 1$. For this purpose we must keep distinct the three cases

$$\begin{cases} (a) \quad 0 < \alpha < 1, \quad |\theta| \le \alpha, \\ (b) \quad 1 < \alpha < 2, \quad |\theta| \le 2 - \alpha, \\ (c) \quad \alpha = 1, \quad |\theta| \le 1. \end{cases}$$

$$(4.11)$$

For $\alpha \neq 1$ the starting point is the Grünwald– Letnikov discretization of fractional derivatives, on which the reader can inform himself from the treatises on fractional calculus, see e.g. [42,46, 50,53], or in the review paper by Gorenflo [18]. However, for our purposes, we must make a clever use of the Grünwald–Letnikov scheme treating separately the two cases (a) and (b) in order to be consistent with a conservative, non-negativity preserving redistribution process. We have

$${}_{h}D_{\pm}^{\alpha}y_{j} = \begin{cases} \frac{1}{h^{\alpha}}\sum_{k=0}^{\infty} (-1)^{k} \binom{\alpha}{k} y_{j\mp k} \\ \text{in the case (a) } 0 < \alpha < 1, \\ \frac{1}{h^{\alpha}}\sum_{k=0}^{\infty} (-1)^{k} \binom{\alpha}{k} y_{j\pm 1\mp k} \\ \text{in the case (b) } 1 < \alpha < 2. \end{cases}$$
(4.12)

Notice the shift of index in the case (b) which among other things has the effect that in the limiting case $\alpha = 2$ (the classical diffusion equation) we obtain the standard symmetric three-point difference scheme. For more details and discussions, see [23]. For $\alpha = 1$ we limit ourselves to consider the case $\theta = 0$ that requires the discretization of the Hilbert transform in (A.11). The hyper-singular integral representation of the symmetric space-fractional derivative (the "Riesz derivative") given by Gorenflo and Mainardi in [25] [as formula (2.20)] readily offers us the discretization

$$_{h}D_{0}^{1}y_{j} = \frac{1}{\pi h}\sum_{k=1}^{\infty}\frac{y_{j+k}-2y_{j}+y_{j-k}}{k(k+1)}.$$
 (4.13)

For $0 < |\theta| \le 1$ we have analogously to (A.10) to take into account the forward/backward difference quotients for ${}_{h}D_{\pm}$,

$$_{h}D_{+} = \frac{y_{j+1} - y_{j}}{h}, \quad _{h}D_{-} = \frac{y_{j} - y_{j-1}}{h}, \quad (4.14)$$

to obtain a non-negative scheme in a proper way. However, we here leave out these "drifts" (in negative or positive directions) from the computation of the transfer coefficients p_k .

Then, inserting the expressions (4.12), (4.13) in the finite-difference equation (4.9), (4.10) yields the *transition law*

$$y_j(t_{n+1}) = \sum_{k=-\infty}^{\infty} p_k y_{j-k}(t_n), \quad j \in \mathbf{Z}, \ \mathbf{n} \in \mathbf{N_0},$$
(4.15)

where the *transfer coefficients* p_k turn out to be, respectively,

$$\begin{cases} p_0 = 1 - \mu(c_+ + c_-) = 1 - \mu \frac{\cos(\theta \pi/2)}{\cos(\alpha \pi/2)}, \\ p_{\pm k} = (-1)^{k+1} \mu \binom{\alpha}{k} c_{\pm}, \quad k = 1, 2, \dots, \\ 0 < \alpha < 1, \ |\theta| \le \alpha, \end{cases}$$
(4.16a)

$$\begin{cases} p_{0} = 1 + \mu \binom{\alpha}{1} (c_{+} + c_{-}) = 1 - \mu \alpha \frac{\cos(\theta \pi/2)}{|\cos(\alpha \pi/2)|}, \\ p_{\pm 1} = -\mu \left[\binom{\alpha}{2} c_{\pm} + c_{\mp} \right], \\ 1 < \alpha < 2, \ |\theta| \leq 2 - \alpha, \\ p_{\pm k} = (-1)^{k} \mu \binom{\alpha}{k+1} c_{\pm}, \quad k = 2, 3, \dots, \end{cases}$$

$$(4.16b)$$

$$\begin{cases} p_0 = 1 - \frac{2\mu}{\pi}, \\ p_{\pm k} = \frac{\mu}{\pi} \frac{1}{k(k+1)}, \quad k = 1, 2, \dots, \end{cases} \quad \alpha = 1, \ \theta = 0, \end{cases}$$
(4.16c)

with the scale parameter

$$\mu := \frac{\tau}{h^{\alpha}}, \quad 0 < \alpha < 2. \tag{4.17}$$

It is straightforward to check in all cases the summation condition $\sum_{k=-\infty}^{\infty} p_k = 1$ Since all $p_{\pm k} \ge 0, k \ne 0$ the non-negativity condition is met if we require $p_0 \ge 0$, i.e. if the scale parameter μ is restricted as follows:

$$0 < \mu := \frac{\tau}{h^{\alpha}} \leqslant \mu_{\max}$$

$$= \begin{cases} \frac{\cos \alpha \pi/2}{\cos \theta \pi/2}, & 0 < \alpha < 1, \ |\theta| \leqslant \alpha, \\ \frac{1}{\alpha} \frac{|\cos \alpha \pi/2|}{\cos \theta \pi/2}, & 1 < \alpha < 2, \ |\theta| \leqslant 2 - \alpha, \\ \frac{\pi}{2}, & \alpha = 1, \ \theta = 0. \end{cases}$$
(4.18)

We note that in both the limits $\alpha \to 1^-$ and $\alpha \to 1^+$ the permissible range of the scaling factor μ is vanishing. In numerical practice the consequence will be that if α is near 1 the convergence is slow: for good approximation we will need a very small step-time τ with respect to the step-length h.⁷

The striking difference to our discretzed Brownian motion is the appearance of arbitrarily large jumps with power-like decay probability for which these discrete models are referred to as *Lévy flights*. See Fig. 3 for a sketch of the transition scheme and Section 5 for the numerical results.

5. The discrete random walk models for the non-Markovian fractional diffusion

Let us now try to generalize the above arguments by adapting them to the non-Markovian cases of our space-time fractional diffusion equa-

⁷ In order to get a continuous transition to the case $\alpha = 1$ we need to consider a different discretization scheme, presented in [19] for the first time and rigorously analyzed in [25]. For this scheme, however, we loose the continuity as $\alpha \rightarrow 2^-$. This means that "there is no free lunch"; we have to pay for the good behaviour at $\alpha = 1$ with bad behaviour at $\alpha = 2$ in a sense described in [19].



Fig. 3. Sketch of the transition scheme for the space-fractional random walker.

tion (2.2). We find it convenient to proceed by steps: we first consider the case of the *strictly time fractional diffusion*, see also [26], and then, by combining these arguments with those for *strictly space fractional diffusion*, we treat the *strictly space-time fractional diffusion*.

5.1. A discrete random walk model for the strictly time fractional diffusion

Discretizing all the variables as for the standard diffusion equation, namely introducing a spacetime mesh of widths h and τ and a discrete variable $y_j(t_n)$ interpreted as in (4.2), the *essential idea* is to replace the *strictly time fractional diffusion* equation by the finite-difference equation

$${}_{\tau}D_{*}^{\beta}y_{j}(t_{n+1}) = \frac{y_{j+1}(t_{n}) - 2y_{j}(t_{n}) + y_{j-1}(t_{n})}{h^{2}},$$

0 < \beta < 1, (5.1)

where the difference operator ${}_{\tau}D_*^{\beta}$ is intended to converge to D_*^{β} as $\tau \to 0$. As usual, we have adopted a symmetric second-order difference quotient in space at level $t = t_n$ for approximating the secondorder space derivative. For ${}_{\tau}D_*^{\beta}$ we require a scheme which must reduce as $\beta = 1$ to a forward difference quotient in time at level $t = t_n$, which is usually adopted for approximating the first-order time derivative. Then, for approximating the time fractional derivative (in Caputo's sense), we adopt a backward Grünwald–Letnikov scheme in time (starting at level $t = t_{n+1}$) which reads

$${}_{\tau}D_{*}^{\beta}y_{j}(t_{n+1}) = \sum_{k=0}^{n+1} (-1)^{k} {\beta \choose k} \frac{y_{j}(t_{n+1-k}) - y_{j}(0)}{\tau^{\beta}},$$

$$0 < \beta < 1.$$
(5.2)

Here, the subtraction of $y_j(0)$ in each term of the sum reflects the subtraction of $f(0^+)$ in formula (B.6) for the Caputo fractional derivative. Combining (5.1) and (5.2), introducing the scaling parameter

$$\mu := \frac{\tau^{\beta}}{h^2}, \quad 0 < \beta \leqslant 1.$$
(5.3)

and using the "empty sum" convention $\sum_{k=p}^{q} = 0$ if q < p (here p = 1 when q = n = 0), we obtain for $n \ge 0$ ($t_0 = 0$):

$$y_{j}(t_{n+1}) = y_{j}(t_{0}) \sum_{k=0}^{n} (-1)^{k} {\binom{\beta}{k}} + \sum_{k=1}^{n} (-1)^{k+1} {\binom{\beta}{k}} y_{j}(t_{n+1-k}) + \mu[y_{j+1}(t_{n}) - 2y_{j}(t_{n}) + y_{j-1}(t_{n})].$$
(5.4)

Thus, (5.4) provides the universal *transition law* from t_n to t_{n+1} valid for all $n \ge 0$. For convenience let us introduce the coefficients c_k , b_m

$$\begin{cases} c_k = (-1)^{k+1} \binom{\beta}{k} = \left| \binom{\beta}{k} \right|, & k \ge 1, \\ b_m = \sum_{k=0}^m (-1)^k \binom{\beta}{k}, & m \ge 0. \end{cases}$$
(5.5)

For $\beta = 1$ (standard diffusion) we note that all these coefficients are vanishing except $b_0 = c_1 = 1$. For $0 < \beta < 1$ they possess the properties

$$\sum_{k=1}^{\infty} c_k = 1, \quad 1 > \beta = c_1 > c_2 > c_3 > \dots \to 0,$$
(5.6)

$$\begin{cases} b_0 = 1 = \sum_{k=1}^{\infty} c_k, & b_m = 1 - \sum_{k=1}^{m} c_k = \sum_{k=m+1}^{\infty} c_k, \\ 1 = b_0 > b_1 > b_2 > b_3 > \dots \to 0. \end{cases}$$
(5.7)

We thus observe that the c_k and the b_m form sequences of positive numbers, not greater than 1, decreasing strictly monotonically to zero. Thanks to the introduction of the above coefficients the universal transition law (5.4) can be written in the following noteworthy form:

$$y_{j}(t_{n+1}) = b_{n}y_{j}(t_{0}) + \sum_{k=1}^{n} c_{k}y_{j}(t_{n+1-k})$$
$$= \mu[y_{j+1}(t_{n}) - 2y_{j}(t_{n}) + y_{j-1}(t_{n})], \qquad (5.8)$$

with the empty sum convention if n = 0. In particular, we get, for n = 0:

$$y_{j}(t_{1}) = (1 - 2\mu)y_{j}(t_{0}) + \mu[y_{j+1}(t_{0}) + y_{j-1}(t_{0})],$$

for $n = 1$:
$$y_{j}(t_{2}) = b_{1}y_{j}(t_{0}) + (c_{1} - 2\mu)y_{j}(t_{1})$$
$$+ \mu[y_{j+1}(t_{1}) + y_{j-1}(t_{1})],$$

for $n \ge 2$:

$$y_j(t_{n+1}) = b_n y_j(t_0) + \sum_{k=2}^n c_k y_j(t_{n+1-k}) + (c_1 - 2\mu) y_j$$
$$\times (t_n) + \mu [y_{j+1}(t_n) + y_{j-1}(t_n)].$$

Observe that $c_1 = \beta$. Scheme (5.8) preserves nonnegativity, if all coefficients are non-negative, hence if

$$0 < \mu = \frac{\tau^{\beta}}{h^2} \leqslant \frac{\beta}{2}.$$
(5.9)

Furthermore it is *conservative*, as we shall prove by induction, i.e.

$$\sum_{j=-\infty}^{+\infty} |y_j(t_0)| < \infty \Rightarrow \sum_{j=-\infty}^{+\infty} y_j(t_n)$$
$$= \sum_{j=-\infty}^{+\infty} y_j(t_0), \quad n \in \mathbf{N}.$$
(5.10)

In fact, putting $S_n = \sum_{j=-\infty}^{+\infty} y_j(t_n)$ for $n \ge 0$, then from (5.8) we get

$$S_{1} = (1 - 2\mu) \sum_{j \neq 1} y_{j}(t_{0}) + \mu \sum_{j \neq 1} y_{j-1}(t_{0}) + \mu \sum_{j \neq 1} y_{j+1}(t_{0}) = S_{0},$$

and for $n \ge 1$ we find always from (5.8), assuming $S_0 = S_1 = \cdots S_n$ already proved,

$$egin{aligned} S_{n+1} &= b_n S_0 + \sum_{k=2}^n c_k S_{n+1-k} + (eta - 2\mu + \mu + \mu) S_n \ &= igg(b_n + \sum_{k=1}^n c_k igg) S_0 = S_0, \end{aligned}$$

using $\beta = c_1$. We have thus proved conservativity. Non-negativity preservation and conservativity mean that our scheme can be interpreted as a *redistribution scheme* of clumps $y_j(t_n)$. For orientation on such aspects and for examples let us quote the works by Gorenflo on conservative difference schemes for diffusion problems, see e.g. [16,17].

The interpretation of our *redistribution* scheme is as follows: the clump $y_j(t_{n+1})$ arises as a weighted-memory average of the (previous) n + 1values $y_j(t_m)$, with m = n, n - 1, ..., 1, 0, with positive weights

$$\beta = c_1, c_2, \dots, c_n, \quad b_n = 1 - \sum_{k=1}^n c_k,$$
 (5.11)

followed by subtraction of $2\mu y_j(t_n)$, which is given in equal parts to the neighbouring points x_{j-1} and x_{j+1} but replaced by the contribution $\mu[y_{j+1}(t_n) + y_{j-1}(t_n)]$ from these neighbouring points. For *random walk* interpretation we consider the $y_j(t_n)$ as probabilities of sojourn at point x_j in instant t_n requiring the normalization condition $\sum_{j=-\infty}^{+\infty} y_j(t_0) = 1$.



Fig. 4. Sketch of the transition scheme for the time-fractional random walker.

For n = 0, Eq. (5.8) means (by appropriate reinterpretation of the spatial index *j*): A particle sitting at x_j in instant t_0 jumps, when *t* proceeds from t_0 to t_1 , with probability μ to the neighbour point x_{j+1} , with probability μ to the neighbour point x_{j-1} , and with probability $1 - 2\mu$ it remains at x_j . For $n \ge 1$, we write (5.8), using $\beta = c_1$, as follows:

$$y_{j}(t_{n+1}) = \left(1 - \sum_{k=1}^{n} c_{k}\right) y_{j}(t_{0}) + c_{n}y_{j}(t_{1}) + c_{n-1}y_{j}(t_{2}) + \dots + c_{2}y_{j}(t_{n-1}) + (c_{1} - 2\mu)y_{j}(t_{n}) + \mu[y_{j+1}(t_{n}) + y_{j-1}(t_{n})].$$
(5.12)

Obviously, all coefficients (probabilities) are non negative, and their sum is 1. But what does it mean? Having a particle, sitting in x_j at instant t_n , where will we find it with which probability at instant t_{n+1} ? From (5.12) we conclude, by re-interpretation of the spatial index *j*, considering the whole history of the particle, i.e. the particle path $\{x(t_0), x(t_1), x(t_2), \ldots, x(t_n)\}$, that if at instant t_n it is in point x_j , there is the contribution $c_1 - 2\mu$ to be again at x_j at instant t_{n+1} , the contribution μ to go to x_{j-1} , the contributions is $c_1 = \beta \leq 1$. So, excluding the case $\beta = 1$ in which we recover the standard diffusion (*Markovian process*), for $\beta < 1$ we have to consider the previous time levels (*non-Markovian process*). Then, from level t_{n-1} we get the contribution c_2 for the probability of staying in x_j also at time t_{n+1} , from level t_{n-2} we get the contribution c_3 for the probability of staying in x_j at time t_{n+1}, \ldots , from level t_1 we get the contribution c_n for the probability of staying in x_j at time t_{n+1} , and finally, from level $t_0 = 0$ we get the contribution b_n for the probability of staying in x_j at time t_{n+1} . Thus, the whole history up to t_n decides probabilistically where the particle will be at instant t_{n+1} .

Let us consider the problem of simulation of transition from time level t_n to t_{n+1} : Assume the particle sitting in x_i at instant t_n . Generate a random number equidistributed in $0 \le \rho < 1$, and subdivide the interval [0,1) as follows. From left to right beginning at zero we put adjacent intervals of length $c_1, c_2, \ldots, c_n, b_n$, for consistency left-closed, right-open. The sum of these is 1. We divide further the first interval (of length c_1) into sub-intervals of length $\mu, c_1 - 2\mu, \mu$. Then we look into which of the above intervals the random number falls. If in first interval with length $c_1 = \mu + \mu$ $(c_1 - 2\mu) + \mu$, then look in which subinterval, and correspondingly move the particle to x_{i-1} , or leave it at x_i or move to x_{i+1} . If the random number falls into one of the intervals with length c_2, c_3, \ldots, c_n (i.e. c_k with $2 \le k \le n$), then move the particle back to its previous position $x(t_{n+1-k})$, which by chance could be identical with $x_j = x(t_n)$. If the random

number falls into the rightmost interval with length b_n then move the particle back to its initial position $x(t_0)$, for which we recommend $x(t_0) = 0$, meaning $y_j(t_0) = \delta_{j0}$, in accordance with the initial condition $u(x, 0) = \delta(x)$ for (2.2) with $\alpha = 2, \theta = 0$.

A sketch of the transition scheme for the random walker is reported in Fig. 4. Besides the diffusive part $(\mu, c_1 - 2\mu, \mu)$ which lets the particle jump at most to neighbouring points, we have for $0 < \beta < 1$ the memory part which gives a tendency to return to former positions even if they are far away. Due to Eqs. (5.6) and (5.7), of course, the probability to return to a far away point gets smaller and smaller the larger the time lapse is from the instant when the particle was there.

5.2. A discrete random walk model for the strictly space-time fractional diffusion

By combining the approach of the preceding Sections 3 and 4.1 we can construct a discrete ran-

dom walk model for the strictly space-time fractional diffusion equation, namely for the case $\{0 < \alpha < 2, 0 < \beta < 1\}$. We replace in (2.2) the (Riesz-Feller) space-fractional derivative by (4.10) and the (Caputo) time-fractional derivative by (5.2). Solving then for the "new" value $y_j(t_{n+1})$ we see that the scaling parameter $\mu = \tau^{\beta}/h^{\alpha}$ plays the essential role in obtaining a scheme with all transition coefficients non-negative. In fact, also here μ must be restricted to a suitable interval $(0, \mu_{max}]$. In analogy to the case of the strictly time fractional diffusion the "discrete" diffusion in space occurs only between the time-level t_n and t_{n+1} and the memory part of the process only straight-backwards in time. However, in contrast to the strictly time fractional diffusion, the discrete diffusion (or the random walker) can now go to any grid point in space, not only to immediate neighbouring grid points. We abstain from presenting here all the lengthy formulas that again need distinction between the three cases: (a) $0 < \alpha < 1$, (b) $1 < \alpha < 2$, and (c) $\alpha = 1$.



Fig. 5. Histogram (left) and a sample path with increments (right) for strictly space fractional diffusion { $\alpha = 1.75, \beta = 1, \theta = 0$ }.



Fig. 6. Histogram (left) and a sample path with increments (right) for strictly space fractional diffusion { $\alpha = 1.75, \beta = 1, \theta = -0.25$ }.



Fig. 7. Histogram (left) and a sample path with increments (right) for strictly space fractional diffusion $\{\alpha = 1.50, \beta = 1, \theta = 0\}$.



Fig. 8. Histogram (left) and a sample path with increments (right) for strictly space fractional diffusion { $\alpha = 1.50, \beta = 1, \theta = -0.50$ }.



Fig. 9. Histogram (left) and a sample path with increments (right) for strictly fractional diffusion { $\alpha = 2, \beta = 0.75, \theta = 0$ }.



Fig. 10. Histogram (left) and a sample path with increments (right) for strictly space fractional diffusion { $\alpha = 2, \beta = 0.50, \theta = 0$ }.



Fig. 11. Histogram (left) and a sample path with increments (right) for strictly space-time fractional diffusion { $\alpha = 1.50, \beta = 0.50, \theta = 0$ }.



Fig. 12. Histogram (left) and a sample path with increments (right) for strictly space-time fractional diffusion { $\alpha = 1.50, \beta = 0.50$, $\theta = -0.50$ }.

6. Numerical results and concluding discussions

Our simulations are all based on 10,000 realizations. In addition to the Brownian motion, see Fig. 2, we have considered a set of eight casestudies, see Figs. 5–12, for our space-time fractional diffusion that, in our opinion, can better

illustrate the state-of-art of this analysis. The sample paths and the corresponding increments are plotted against the time steps up to 500, see the right figure-plates, while the histograms refer to densities at t = 1 for $|x| \le 5$, see the left figure-plates. All the plots were drown by using the MATLAB system. The relevant parameters

α	θ	β	$\mu_{ m max}$	$h_{ m S}$	$ au_{ m S}$	$h_{ m H}$	$ au_{ m H}$
2	0	1	0.50	$7.0 imes10^{-2}$	$2.0 imes 10^{-3}$	0.20	$2.5 imes 10^{-2}$
1.75	0	1	0.53	$4.7 imes 10^{-2}$	$2.0 imes 10^{-3}$	0.33	$2.0 imes 10^{-2}$
1.75	25	1	0.57	$4.5 imes10^{-2}$	$2.0 imes10^{-3}$	0.33	$2.0 imes10^{-2}$
1.50	0	1	0.47	$3.0 imes 10^{-2}$	$2.0 imes 10^{-3}$	0.20	$1.0 imes10^{-2}$
1.50	50	1	0.67	$2.4 imes 10^{-2}$	$2.0 imes 10^{-3}$	0.20	$1.0 imes10^{-2}$
2	0	0.75	0.37	0.17	$2.0 imes 10^{-3}$	0.25	$5.0 imes 10^{-3}$
2	0	0.50	0.25	0.47	$2.0 imes 10^{-3}$	0.50	$2.5 imes 10^{-2}$
1.50	0	0.50	0.24	0.38	$2.0 imes 10^{-3}$	0.50	$5.0 imes 10^{-3}$
1.50	50	0.50	0.33	0.30	$2.0 imes 10^{-3}$	0.50	$5.0 imes 10^{-3}$

Table 1 The relevant parameters for the simulations

 $\alpha =$ space-fractional order, $\theta =$ skewness, $\beta =$ time-fractional order, $\mu = \tau^{\beta}/h^{\alpha} =$ scaling parameter, $h_{\rm S} =$ space-step, $\tau_{\rm S} =$ time-step for sample paths, $h_{\rm H} =$ space-step, $\tau_{\rm H} =$ time-step for histograms.

 $\alpha, \theta, \beta, \mu, h$ and τ used in the figures are reported in Table 1. For convenience we have also reported the maximum value of the scaling parameter μ as can be deduced from our theoretical analysis. More details can be found in [45].

In practice, in our numerical studies there is required truncation for two different causes. As in the classical Brownian motion, a trivial truncation is required if a priori one wants a definite region of space to be considered in which the walk takes place. Then, particles leaving this space have been ignored. However, if $0 < \alpha < 2$, at variance with our discretized Brownian motion, we now have an infinite number of transition probabilities. Since it is impossible to simulate all infinitely many discrete probabilities, so the size of possible jumps must be limited to a maximal possible jump length.

Figs. 5-8 are concerning four cases (both symmetrical and extremal) of strictly space fractional *diffusion*: { $\alpha = 1.75, \beta = 1$ } with $\theta = 0, -0.25$ and $\{\alpha = 1.5, \beta = 1\}$ with $\theta = 0, -0.5$. From the sample paths one can recognize the "wild" character (with large jumps) of the Lévy flights with respect to the "tame" character of the Brownian motion outlined in Fig. 2. Figs. 9 and 10 are concerning two cases of strictly time fractional diffusion: $\{\alpha = 2, \beta = 0.75\}$ and $\{\alpha = 2, \beta = 0.50\}$; here the paths exhibit the memory effect visible in a kind of stickiness combined with occasional jumps to points previously occupied, in distinct contrast to the rather tame behaviour in case $\{\alpha = 2, \beta = 1\}$ (simulation of Brownian motion). Finally, Figs. 11 and 12 are concerning two cases of strictly spacetime fractional diffusion: { $\alpha = 1.50, \beta = 0.50, \theta =$ 0} and { $\alpha = 1.50, \beta = 0.50, \theta = -0.50$ }, where the combined effects of the previous cases are present.

We have used our discrete models for simulation of particle trajectories by interpreting our redistribution schemes as descriptions of Markov chains with infinitely many states, namely the possible positions x_j . However, as they have the form of difference schemes in a regular space-time grid they can "in principle" also serve for the purpose of approximate computation of the temporal evolution of the density u(x, t). Namely, we are expecting that $u(x_i, t_n)$ is approximated by $y_j(t_n)/h$. In fact, from their essential properties (conservativity and preservation of non-negativity, see (4.6), (5.11) and (5.12)) it can by standard methods of numerical analysis been shown that our models interpreted as difference schemes are stable and consistent, hence convergent. We say "in principle" because for practical application appropriate truncations are required. That there is also convergence in the chain interpretation (complete convergence in the stochastic sense) can be shown by methods of Fourier and Laplace analysis, see [24,25] for the space-fractional diffusion.

There are other methods of simulating spacefractional diffusion processes. Without attempting to be exhaustive, without going into details and without attempting to give a survey of the existing and growing literature on the subject let us mention a few possibilities: generations of random numbers distributed according to a given stable law, random walks discrete in time but continuous in space (namely jumps to arbitrary places in space, not only to grid points), let us mention here the Chechkin-Gonchar random walk (suggested and used in [9] and mathematically analyzed in [25]), and simulation via compound Poisson processes (with proper scaling of space and time), see [27]. In the special field of methods, discrete in space and time, there also are other methods available (see the methods sketched in [19] and discussed in detail in [24,25] of which, in particular, the Gillis-Weiss random walk must be cited, see [14]). Final word: Still many questions are open in this challenging and fascinating field of research!

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Appendix A. The Riesz–Feller space fractional derivatives

In this section we provide the explicit expression of the *Riesz–Feller* fractional derivative ${}_{x}D^{\alpha}_{\theta}$ which, according to (2.4) is defined as the pseudo-differential operator with symbol

$$\widehat{{}_{x}D^{\alpha}_{\theta}} = -|\kappa|^{\alpha} \mathrm{e}^{\mathrm{i}(\mathrm{sign}\,\kappa)\theta\pi/2}. \tag{A.1}$$

Let us now express more properly the operator $_{x}D_{\theta}^{\alpha}$ as inverse of a suitable integral operator $_{x}I_{\theta}^{\alpha}$, whose symbol is required to be $|\kappa|^{-\alpha}e^{-i(\operatorname{sign}\kappa)\theta\pi/2}$, so we may write

$$_{x}D^{\alpha}_{\theta} := -_{x}I^{-\alpha}_{\theta}.$$
 (A.2)

This integral operator was found by Feller [12] in 1952 generalizing the approach by Marcel Riesz to Fractional Calculus, see e.g. [51], and it is referred to as *Feller potential* by Samko et al. [53].⁸ Using our notation, the *Feller potential* reads

$${}_{x}I^{\alpha}_{\theta}f(x) = c_{-}(\alpha,\theta)_{x}I^{\alpha}_{+}f(x) + c_{+}(\alpha,\theta)_{x}I^{\alpha}_{-}f(x), \quad (A.3)$$

where, if $0 < \alpha < 2$, $\alpha \neq 1$,

$$c_{+}(\alpha, \theta) = \frac{\sin[(\alpha - \theta)\pi/2]}{\sin(\alpha\pi)},$$

$$c_{-}(\alpha, \theta) = \frac{\sin[(\alpha + \theta)\pi/2]}{\sin(\alpha\pi)},$$
(A.4)

and, by passing to the limit (with $\theta = 0$) $c_+(2,0) = c_-(2,0) = -1/2$. In (A.4) the operators ${}_xI^{\alpha}_{\pm}$ denote the *Riemann–Liouville fractional integrals*, also known as *Weyl fractional integrals*), defined as

$$\begin{cases} {}_{x}I^{\alpha}_{+}f(x) = \frac{1}{\Gamma(\alpha)} \int_{-\infty}^{x} (x-\xi)^{\alpha-1}f(\xi) \, \mathrm{d}\xi, \\ {}_{x}I^{\alpha}_{-}f(x) = \frac{1}{\Gamma(\alpha)} \int_{x}^{+\infty} (\xi-x)^{\alpha-1}f(\xi) \, \mathrm{d}\xi. \end{cases}$$
(A.5)

We note that whereas the coefficients c_{\pm} can loose their meaning when α is an integer, the Riemann–Liouville integral operators I_{\pm}^{α} are well defined in their action on rapidly decreasing in-

$$\left(|\kappa|\mathrm{e}^{-\mathrm{i}(\mathrm{sign}\,\kappa)\delta}\right)^{-\alpha}$$
 so $\delta = -\frac{\pi}{2}\frac{\theta}{\alpha}, \ \theta = -\frac{2}{\pi}\alpha\delta$

tegrable functions, for any $\alpha \ge 0$, being set equal to the identity operator when $\alpha = 0$, for convenience (justified by passage to the limit $\alpha \to 0$). In the particular case $\theta = 0$ we get

$$c_{+}(\alpha, 0) = c_{-}(\alpha, 0) = \frac{1}{2\cos(\pi\alpha/2)},$$
 (A.6)

and thus we recover the *Riesz potential*, see e.g. [53],

$${}_{x}I_{0}^{\alpha}f(x) := \frac{1}{2\Gamma(\alpha)\cos(\pi\alpha/2)} \int_{-\infty}^{+\infty} |x - \xi|^{\alpha - 1} f(\xi) \, \mathrm{d}\xi.$$
(A.7)

The Riesz and the Feller potentials are well defined if the index is located in the interval (0, 1), or in the interval (0, 2), and we have the semigroup property, ${}_{x}I^{\alpha}_{\theta} {}_{x}I^{\beta}_{\theta} = {}_{x}I^{\alpha+\beta}_{\theta}$, if $0 < \alpha < 1$, $0 < \beta < 1$ and $\alpha + \beta < 1$. Then, following Feller, we define by analytic continuation the pseudo-differential operator (A.2) in the whole range $0 < \alpha \leq 2$, $\alpha \neq 1$ as

$${}_{x}D^{\alpha}_{\theta} := -[c_{+}(\alpha,\theta)_{x}D^{\alpha}_{+} + c_{-}(\alpha,\theta)_{x}D^{\alpha}_{-}],$$

$$0 < \alpha \leqslant 2,$$
(A.8)

where ${}_{x}D^{\alpha}_{+} := {}_{x}I^{-\alpha}_{+}$ and ${}_{x}D^{\alpha}_{-} := {}_{x}I^{-\alpha}_{-}$ are the inverses of the operators ${}_{x}I^{\alpha}_{+}$ and ${}_{x}I^{\alpha}_{-}$, respectively, and are referred to as the *Weyl fractional derivatives*. When $\theta = 0$ the *Riesz–Feller derivative* can be simply referred to as *Riesz derivative*. We note from (A.4) the property $c_{\pm}(\alpha, \theta) = c_{\mp}(-\alpha)$. For integral representations of the operators ${}_{x}I^{-\alpha}_{+}$ see [53]; we have

$$_{x}I_{\pm}^{-\alpha} = \begin{cases} \pm \frac{\mathrm{d}}{\mathrm{d}x} \begin{pmatrix} xI_{\pm}^{1-\alpha} \end{pmatrix} & \text{if } 0 < \alpha \leqslant 1, \\ \frac{\mathrm{d}^{2}}{\mathrm{d}x^{2}} \begin{pmatrix} xI_{\pm}^{2-\alpha} \end{pmatrix} & \text{if } 1 < \alpha \leqslant 2. \end{cases}$$
(A.9)

For $\alpha = 2$ ($\theta = 0$) we recover the standard second derivative. In fact, from (A.6) $c_+(2,0) = c_-(2,0) = -1/2$, so from (A.8), (A.9)

$$_{x}D_{0}^{2} = -I_{0}^{-2} = (_{x}I_{+}^{-2} + _{x}I_{-}^{-2})/2 = \left(\frac{\mathrm{d}^{2}}{\mathrm{d}x^{2}} + \frac{\mathrm{d}^{2}}{\mathrm{d}x^{2}}\right)/2$$

= $\frac{\mathrm{d}^{2}}{\mathrm{d}x^{2}} = _{x}D^{2}.$

For $\alpha = 1(|\theta| \le 1)$ we have (from calculation with symbols)

$$_{x}D_{\theta}^{1}f(x) = [\cos(\theta\pi/2)_{x}D_{0}^{1} + \sin(\theta\pi/2)_{x}D]f(x),$$
(A.10)

⁸ We must note that in his original paper Feller used a skewness parameter δ different from our θ ; the symbol of the Feller potential is

Feller and Samko, Kilbas and Marichev thus use I^{α}_{δ} where their δ is related to our θ as above.

where
$$_{x}D_{0}^{1}f(x) = -_{x}D[_{x}Hf(x)]$$
, with

$${}_{x}Hf(x) = \frac{1}{\pi} \int_{-\infty}^{+\infty} \frac{f(\xi)}{x - \xi} d\xi$$
$$= \frac{1}{\pi} \int_{-\infty}^{+\infty} \frac{f(x - \xi)}{\xi} d\xi.$$
(A.11)

In (A.11) $_{x}H$ denotes the Hilbert transform (with symbol $_{x}H(\kappa) = i \operatorname{sign} \kappa$) and its singular integral is understood in the Cauchy principal value sense. We note that in the limiting extremal cases $\theta = \pm 1$ we recover the standard first derivative, i.e. $_{x}D_{\pm 1}^{1} = \pm_{x}D$.

Appendix B. The Riemann–Liouville and Caputo time fractional derivatives

The purpose of this section is to clarify for the interested reader the main differences between the *Caputo* fractional derivative ${}_{t}D_{*}^{\beta}$ adopted in this paper, see (2.8), and the *Riemann–Liouville* fractional derivative ${}_{t}D^{\beta}$ usually adopted in the literature. Any formula here is valid for t > 0, with the assumption that the function f(t) has a finite limit as $t \rightarrow 0^{+}$. We recall

$$_{t}D^{\beta}f(t) := \begin{cases} \frac{1}{\Gamma(1-\beta)} \frac{\mathrm{d}}{\mathrm{d}t} \int_{0}^{t} \frac{f(\tau)}{(t-\tau)^{\beta}} \mathrm{d}\tau, & 0 < \beta < 1, \\ \frac{\mathrm{d}f(t)}{\mathrm{d}t}, & \beta = 1. \end{cases}$$
(B.1)

The two fractional derivatives are related to the Riemann–Liouville fractional integral as follows. The Riemann–Liouville fractional integral is

$${}_{r}J^{\mu}f(t) := \frac{1}{\Gamma(\mu)} \int_{0}^{t} f(\tau)(t-\tau)^{\mu-1} \,\mathrm{d}\tau, \quad \mu > 0$$
(B.2)

(with the convention ${}_{t}J^{0}f(t) = f(t)$) and is known to satisfy the semigroup property ${}_{t}J^{\mu}{}_{t}J^{\nu} = {}_{t}J^{\mu+\nu}$, with $\mu, \nu > 0$. For any $\beta > 0$ the Riemann–Liouville fractional derivative is defined as the *left* inverse of the corresponding fractional integral (like the derivative of any integer order), namely ${}_{t}D^{\beta}{}_{t}J^{\beta}f(t)$ = f(t). Then for $\beta \in (0, 1]$, we have

$${}_{t}D^{\beta}f(t) := {}_{t}D^{1}{}_{t}J^{1-\beta}f(t),$$

$${}_{t}D^{\beta}_{*}f(t) := {}_{t}J^{1-\beta}{}_{t}D^{1}f(t),$$
 (B.3)

$${}_{\prime}J^{\beta}{}_{\iota}D^{\beta}_{*}f(t) = {}_{\iota}J^{\beta}{}_{\iota}J^{1-\beta}{}_{\iota}D^{1}f(t) = {}_{\iota}J^{1}{}_{\iota}D^{1}f(t)$$

= $f(t) - f(0^{+}).$ (B.4)

Recalling the rule

$${}_{t}D^{\beta}t^{\gamma} = \frac{\Gamma(\gamma+1)}{\Gamma(\gamma+1-\beta)}t^{\gamma-\beta}, \quad \beta \ge 0, \ \gamma > -1,$$
(B.5)

it turns out for $0 < \beta \leq 1$,

$${}_{t}D_{*}^{\beta}f(t) = {}_{t}D^{\beta}[f(t) - f(0^{+})]$$

= ${}_{t}D^{\beta}f(t) - f(0^{+})\frac{t^{-\beta}}{\Gamma(1-\beta)}.$ (B.6)

Note that for $\beta = 1$ the two types of fractional derivative coincide.

The *Caputo* fractional derivative represents a sort of regularization in the time origin for the *Riemann–Liouville* fractional derivative and satisfies the relevant property of being zero when applied to a constant. For more details on this fractional derivative (and its extension to higher orders) we refer the interested reader to Gorenflo and Mainardi [22], Podlubny [50] and Butzer and Westphal [2].

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